

Buffer gas induced collision shift for the $^{88}\text{Sr } ^1S_0 - ^3P_1$ clock transition

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Precision saturation spectroscopy of the $^{88}\text{Sr } ^1S_0 - ^3P_1$ is performed in a vapor cell filled with various rare gas including He, Ne, Ar, and Xe. By continuously calibrating the absolute frequency of the probe laser, buffer gas induced collision shifts of $\sim\text{kHz}$ are detected with gas pressure of 1-20 mTorr. Helium gave the largest fractional shift of $1.6 \times 10^{-9} \text{ Torr}^{-1}$. Comparing with a simple impact calculation and a Doppler-limited experiment of Holtgrave and Wolf [Phys. Rev. A **72**, 012711 (2005)], our results show larger broadening and smaller shifting coefficient, indicating effective atomic loss due to velocity changing collisions. The applicability of the result to the $^1S_0 - ^3P_0$ optical lattice clock transition is also discussed.

The latest progress of optical clocks is so rapid that their capability is surpassing that of microwave clocks. A fractional accuracy of $\sim 10^{-17}$ was demonstrated with single-ion systems [1]. This remarkable precision, on the other hand, requires us to consider various systematic shifts which have been so far ignored. Neutral atomic clocks, particularly fountain clocks, have been often limited in their accuracy by binary collisions [2]. Besides that, whether the clock is neutral-atom or ionic system, collision shifts induced by background gas (BG) [3] may ultimately limit the accuracy in this unprecedented level. However, the characterization of BG induced collision shifts is difficult in real clocks because the deliberate injection of buffer gas for the shifting of atomic resonance normally causes loss of atoms, limiting the coherence time needed to evaluate subtle BG induced shifts.

From the beginning of laser spectroscopy, frequency shifts and broadening caused by background gas collisions have been investigated both experimentally and theoretically [3]. The spin-forbidden transitions $^1S_0 - ^3P_1$ of atoms with two outer electrons have been major testing grounds due to their fairly narrow natural linewidths and level simplicity free from hyperfine structure. Experiments were done for Ca [4], Sr [4, 5, 6], Cd [7, 8], Hg [9, 10] and Yb [11] with perturber pressures of $10^0 - 10^2$ Torr. In these experiments, shifts and broadening exceeds 10 MHz, and high order effects such as collision time asymmetry or speed-dependence of collisional broadening and shifting can be observed [12, 13]. The higher-order effects, however, disturb the accurate measurement of the linear collision shift and broadening, which is necessary to evaluate systematic shifts of atomic clocks.

In this paper, buffer gas induced collision shifts for the $^{88}\text{Sr } ^1S_0 - ^3P_1$ transition is reported based on pre-

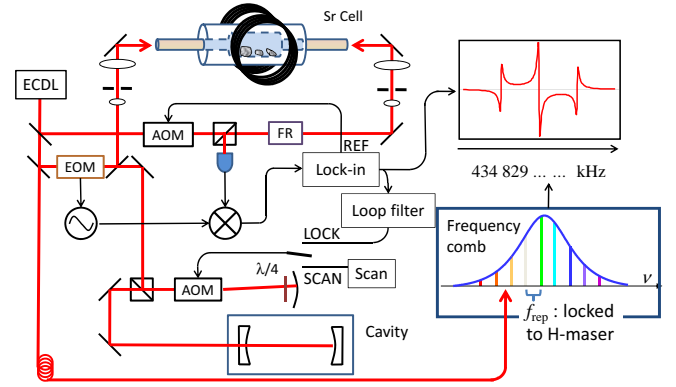


FIG. 1: (Color online) Experimental configuration of the precision saturated absorption spectroscopy. ECDL: extended cavity diode laser, FR: Faraday rotator, EOM: Electro-optic modulator, AOM: acousto-optic modulator

cision saturation absorption spectroscopy with the FM sideband technique [14, 15]. Continuously monitoring the probe laser frequency using a frequency comb referenced to International Atomic Time (TAI), we compared the resonance frequency with the true atomic resonance measured with ultra-cold atoms [16]. The detected background gas induced collision shift was $\sim\text{kHz}$, which is, to the best of our knowledge, the finest definitive measurement of optical frequency shifts induced by foreign-gas collisions.

The $(5s^2)^1S_0 - (5s5p)^3P_1$ transition at $\lambda = 689 \text{ nm}$ has a natural line width $\gamma = 2\pi \times 7.4 \text{ kHz}$ which corresponds to the saturation intensity $I_s = 2\pi^2\hbar c\gamma/3\lambda^3 = 2.9 \mu\text{W}/\text{cm}^2$, where c and $2\pi\hbar$ is the speed of light and the Planck constant, respectively. The experimental setup of our saturated absorption spectroscopy is shown in Fig. 1. The spectral width of an extended cavity diode laser (ECDL) is reduced to less than 500 Hz by locking the laser frequency to a cavity with a basic Pound-

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Drever-Hall (PDH) method [17]. The modulation index and frequency ω_M are 0.17 and $2\pi \times 21.7$ MHz, respectively. This modulation is used for the saturation absorption spectroscopy as well. Strontium is contained in a heat pipe with Brewster window plates at both ends. The typical temperature of the pipe is 728 K. The two ends of the active area (20 cm length) are water-cooled so that the strontium is not deposited on the windows. The buffer gas is introduced from one end of the pipe, and Pirani gauges are installed at both ends for the confirmation of uniform pressures in the tube. The magnetic sublevels are well defined by a bias magnetic field of 3 G produced by a Helmholtz coil pair. Pump and probe beam of the saturation spectroscopy have the same linear polarization which excites the π transition of $\Delta m_J = 0$. Pinholes are placed in both arms to overlap the pump and probe beam. The peak intensity and $1/e^2$ radius of the pump beam (probe beam) are $270 \mu\text{W}/\text{cm}^2$ and 4.2 mm ($50 \mu\text{W}/\text{cm}^2$ and 3.3 mm), respectively. After interacting with Sr vapor, the polarization of the probe beam is rotated 90 degrees and coupled out by a polarization beam splitter. The pump beam is frequency shifted by 80 MHz from the probe beam in order to avoid DC noise caused by the interference between the two beams. In addition, the pump beam is chopped with a frequency of 5 kHz to extract the saturation signal induced by the pump beam. The signal is first demodulated at ω_M and then lock-in detected. The absolute frequency of the laser is always monitored by a frequency comb referenced to International Atomic Time (TAI).

A sample of PDH phase detection signal with helium buffer gas of 4×10^{-4} Torr is shown in Fig. 2(a). The absolute frequency of the probe laser is expressed as a differential frequency $(f_{\text{pump}} + f_{\text{probe}})/2 - f_0$, where f_{pump} and f_{probe} is pump and probe laser frequency, respectively, and the true atomic resonance f_0 is the frequency precisely measured in [16] from ballistically expanding ultracold atoms. The full width half maximum (FWHM) is nonlinear-fitted to be 107 kHz, which agrees with saturation broadening (80 kHz) and transit time broadening (31 kHz) expected from the beam intensity and diameter. This dispersion shape works as an error signal to lock the laser. The stability of this Sr vapor cell clock is shown in Fig. 2(b) as modified Allan deviation against a Hydrogen maser. The bottom of the instability is 4×10^{-14} at 100 s. The stability is limited by fluctuations of buffer gas and vapor pressure or pump beam intensity that changes the light-pressure-induced lineshape asymmetry [18].

We measured the pressure dependence of the background gas induced broadening and shift for various background gases. The buffer gases are rare gases, namely helium, neon, argon, and xenon. The attempts to fill with nitrogen failed since the gas was quickly adsorbed by solid strontium. The broadening is determined from the dispersion curve as in Fig. 2(a). Locking the laser frequency to the zero-crossing of dispersion curve, the resonance is determined by a frequency comb that is referenced to TAI. We determined the slope of the colli-

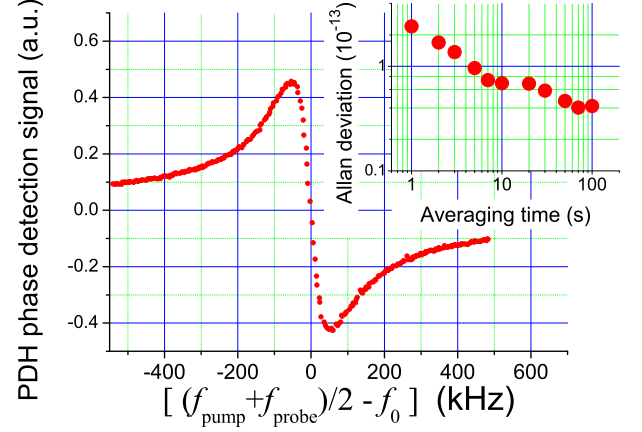


FIG. 2: (a) Dispersion signal of FM side band spectroscopy. The offset frequency of $f_0 = 434\,829\,121\,312$ kHz is the true atomic resonance measured in [16]. (b) Modified Allan standard deviation of the 689 nm laser frequency tightly locked to the resonance signal obtained by the saturation spectroscopy.

sional broadening and the frequency shift as a function of the buffer gas pressure by doing a least square fit, and the result is summarized in Table I.

The frequency shift data are summarized in Fig. 3. The experimental condition described above has resulted in a residual offset of -4 kHz from the true atomic resonance. Part of the reason is a shift caused by homonuclear two body collisions. By changing the temperature of the heat pipe, we measured the Sr-Sr collision shift coefficient to be $-1.3 \times 10^{-10} \text{ Hz cm}^3$ [19]. This effect causes a shift of -1 kHz for the condition of Fig. 3. We suspect that the most probable reason for the remaining -3 kHz is residual Doppler shifts. The spatial filters in the two arms still permit a mutual beam tilt of $\theta = 62 \mu\text{rad}$. In the case that atoms are directed as in an atomic beam, this tilt causes Doppler shift of $k_L v_m \sin(\theta/2) = 15 \text{ kHz}$ maximum, where k_L is the wave number of the light, $v_m = \sqrt{2k_B T/m} = 371 \text{ m/s}$ is most probable speed of Sr atoms at gas temperature $T = 455^\circ\text{C}$, m is the Sr mass and k_B is the Boltzmann constant. Our spectroscopy is ideally performed with an isotropic velocity distribution of atoms. However, a slightly biased velocity distribution may be produced in the heat pipe due to an inhomogeneous temperature distribution of the pipe. Since the beam passes above bulk solid strontium, the distribution could be biased upward. Other possibilities are light-pressure-induced lineshape asymmetry [18] and re-coils due to photon absorption and emission [20, 21, 22].

Only the collisions with He show a positive shift coefficient, whereas Ne, Ar, Xe show negative coefficients. The open symbols at lower pressures are excluded from the fitting since the vapor pressure of strontium is not small enough comparing with the buffer gas pressure. Even slightly above the vapor pressure, the buffer gas pressure wasn't uniform at pressures below 3×10^{-3} Torr.

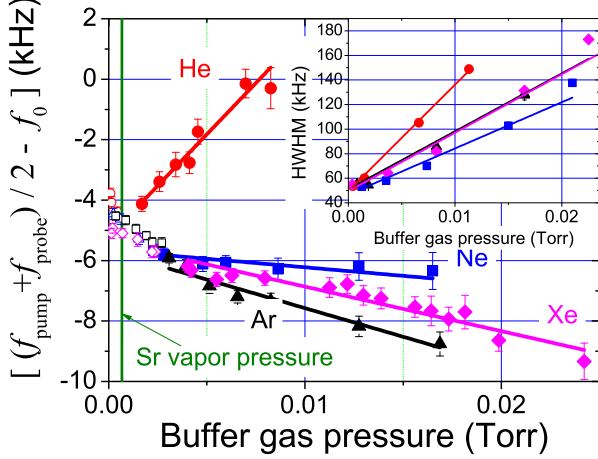


FIG. 3: (Color online) Shifts of the resonance frequency induced by various kinds of buffer gas. The corner of the slope around 3 mTorr is attributed to the inhomogeneous pressure distribution in the cell, which occurs due to insufficient buffer gas pressure relative to strontium vapor. Inset is the broadening of the Doppler-free atomic resonance.

We added the buffer gas from one side of the pipe, and the spatial gradient of the gas pressure was observed by Pirani gauges placed at the both ends of the heat pipe.

The collisional width Γ (HWHM) and shift Δ in the impact approximation are proportional to the density of perturbors N and are given by the well-known expression [3, 23]:

$$\Gamma + i\Delta = N \int d^3\vec{v}_r f_\mu(\vec{v}_r) v_r \int_0^\infty d\rho \rho \left\{ 1 - \left\langle S_{ii} S_{ff}^{-1} \right\rangle_{Av.} \right\} \quad (1)$$

where $f_\mu(\vec{v}_r)$ is the Maxwellian distribution of relative velocities \vec{v}_r of colliding absorber and perturber having reduced mass μ , and ρ is the impact parameter. In this work we report results obtained in the mean speed approximation in which calculations are done just for the mean relative speed $\bar{v} = \sqrt{8k_B T / (\pi\mu)}$ of colliding atoms. For the intercombination transition $1S_0 - 3P_1$, in the adiabatic approximation the angular average of scattering matrix elements in the initial and final states $\left\langle S_{ii} S_{ff}^{-1} \right\rangle_{Av.} = \frac{1}{3} e^{-i\eta_{A-X}} + \frac{2}{3} e^{-i\eta_{B-X}}$ can be expressed in terms of phase shifts η_{A-X} and η_{B-X} induced by the quasimolecular transitions $A^30^+ - X^10^+$ and $B^31 - X^10^+$, respectively [13, 24]. These phase shifts are integrals $\eta = \int_{-\infty}^{+\infty} dt \Delta V(r(t)) / \hbar$ of the excited and ground electronic state potentials differences $V(A^30^+) - V(X^10^+)$ and $V(B^31) - V(X^10^+)$ for Sr-Rg(He, Ne, Ar, Xe) systems. For simplicity we assumed straight line trajectories with the interatomic separation $r(t) = \sqrt{\rho^2 + v_r^2 t^2}$.

The interaction potentials are approximated here by a modified Lennard-Jones formula $V(r) = C_{12}/r^{12} + C_8/r^8 + C_6/r^6$. Following Hindmarsh et al. [25] C_{12} coefficients are estimated using Hindmarsh radii $r_H(\text{Rg})$

for rare gases Rg [26] and $r_H(\text{Sr})$ for Sr. The Hindmarsh radii $r_H(\text{Sr})$ have been estimated using the charge densities calculated by Mitroy and Zhang [27]. They were found to be $9.4773 a_0$ and $10.8801 a_0$ in the ground $1S_0$ and excited $3P_{0,1,2}$ electronic state of Sr atom, respectively. The atomic unit of length is $a_0 \approx 0.05292 \text{ nm}$. The dispersion coefficients C_6 and C_8 for Sr-Rg were also calculated by Mitroy and Zhang [27]. The dispersion coefficient given for excited states $3\Sigma^+$ and 3Π were converted to appropriate coefficients for states A^30^+ and B^31 using the following asymptotic relations: $C_6(A^30^+) = C_6(3\Pi)$ and $C_6(B^31) = \frac{1}{2}C_6(3\Sigma^+) + \frac{1}{2}C_6(3\Pi)$. The term C_8/r^8 included in our calculations allows for a better description of the interaction anisotropy in the excited state of Sr. We note that the addition of the C_8/r^8 and C_{10}/r^{10} terms to the interaction model changes the calculated broadening about a factor of two. The change can be even bigger in the case of the shift and for Ne can lead to a change of its sign. We also verified that if the thermal average is neglected, the resulting error is less than 5%. The positive collisional shift observed for He is caused by the dominating influence of the repulsive interaction C_{12}/r^{12} on the collisional broadening and shifting.

Table I shows the comparison of the pressure broadening and shifting coefficients measured in this work with those calculated in the impact approximation, Eq. (1), and experimental results obtained by other authors. There is only one full set of data for rare gas pressure broadening and shifting of strontium intercombination line reported by Holtgrave and Wolf [6]. Results of Ref. [6] were obtained in the Doppler-limited regime at a temperature of 660(47) K, where the magnitude of the broadening and shifts exceed 100 MHz. Note that collisional broadening cross section determined by Holtgrave and Wolf agree very well with those obtained by Crane et al [4] for Ar and differ by 23% from those for Ne. In Ref. [4] also the Doppler-limited spectroscopical method was used. It is clear that our results do not agree with those obtained using Doppler limited spectroscopy [4, 6]. On the other hand our results agree well with those determined by Tino *et al.* [5] for Ar in the Doppler free regime, see Table I.

The comparison of calculated results seems to give a bit better agreement with results obtained in the Doppler-limited case. Clearly Doppler-free results are characterized by bigger collision broadening coefficients and smaller magnitude of collision-shifting coefficients than those obtained from the Doppler-limited case. This can be attributed to velocity changing collisions [28] which change the absorber velocity significantly enough to push away the absorber from the condition of Doppler-free resonance. Therefore, such events of collisions which can cause large frequency shifts in the Doppler-limited regime would not contribute to the signal in the Doppler-free case. Being a loss mechanism, it will lead to broadening but not to shifting of the line. This effect needs further theoretical and experimental investigation. More-

TABLE I: The comparison of our experimental pressure broadening and shifting coefficients with those calculated in the impact approximation, as well as with experimental results obtained by Holtgrave and Wolf [6] in the Doppler limited regime and obtained by Tino *et al.* [5] in the Doppler free regime. All data are given in MHz/Torr and the one standard deviation uncertainties are given in parentheses.

Perturber	Doppler limited				Doppler free			
	Experiment		Impact calculation		Experiment		Impact calculation	
	Holtgrave and Wolf		($T = 660$ K)		This work		($T = 728$ K)	
	Γ/p	Δ/p	Γ/p	Δ/p	Γ/p	Δ/p	Γ/p	Δ/p
He	3.79(17)	+1.23(6)	3.89	+0.68	8.67(10)	+0.68(8)	3.68	+0.66
Ne	2.60(13)	-0.64(4)	1.17	-0.31	3.76(5)	-0.06(3)	1.10	-0.28
Ar	2.45(4)	-1.45(4)	3.30	-1.54	4.76(8)	-0.22(2)	3.05	-1.47
					5.5(5) ^a			
Xe	3.44(23)	-1.51(3)	3.35	-1.36	4.71(6)	-0.12(1)	3.17	-1.33

^aResult obtained by Tino *et al.* [5].

over in these investigations it would be very useful to carry out proper close-coupling broadening and shifting calculations [29] based on *ab initio* potentials.

The cold strontium system is currently studied extensively in many national standards institutes due to its application to optical lattice clocks [30]. For the clock transition $^1S_0 - ^3P_0$, the interaction of the ground state rare gas atom with the excited clock atom at 3P_0 state is described by just one potential curve denoted $^30^-$ approaching dissociation limit. The dispersion coefficient describing this interaction fulfills the following relation $C_6(^30^-) = C_6(^3\Pi)$ and with a good approximation is equal to $C_6(A^30^+)$ investigated here. Moreover, in the case of collisions of clock atoms with background gas atoms, the velocity-changing collisions will also play a crucial role leading to relatively large collisional broadening and small collisional shifting. Therefore, results

obtained here as well as the discrepancy of the shift coefficients depending on Doppler-free or -limited system should help to correctly estimate the impact of residual gas induced collision shift on optical atomic clocks. The values of fractional shift coefficients $< 2 \times 10^{-9} \text{Torr}^{-1}$ obtained here can be treated as an upper limit for shifts caused by the residual gas in optical lattice clocks which in fact should be even smaller.

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- [1] T. Rosenband, *et al.*, Science **319**, 1808 (2008).
 - [2] Y. Sortais, *et al.*, Phys. Rev. Lett. **85**, 003117 (2000).
 - [3] N. Allard and J. Kielkopf, Rev. Mod. Phys. **54**, 1103 (1982); J. Szudy and W. E. Baylis, Phys. Rep. 266, 127 (1996).
 - [4] J. K. Crane, *et al.*, Phys. Rev. A **49**, 1666 (1994).
 - [5] G. M. Tino, *et al.*, Appl. Phys. B **55**, 397 (1992).
 - [6] J. C. Holtgrave and P. J. Wolf, Phys. Rev. A **72**, 012711 (2005).
 - [7] K. J. Dietz, *et al.*, J. Phys. B **13**, 2749 (1980).
 - [8] A. Bielski, *et al.*, Acta. Phys. Pol. B **33**, 2267 (2002); R. S. Trawinski, *et al.*, Phys. Rev. A **74**, 022716 (2006).
 - [9] J. P. Jacobs and R. B. Warrington, Phys. Rev. A **68**, 032722 (2003).
 - [10] T. N. Anderson, *et al.*, Appl. Phys. B **87**, 341 (2007).
 - [11] D. F. Kimball, *et al.*, Phys. Rev. A **60**, 1103 (1999).
 - [12] I. Shannon, *et al.*, J. Phys. B **19**, 1409 (1986).
 - [13] A. Bielski, *et al.*, Phys. Rev. A **62**, 032511 (2000).
 - [14] J. L. Hall, *et al.*, Appl. Phys. Lett. **39**, 680 (1981).
 - [15] G. C. Bjorklund, *et al.*, Appl. Phys. B **32**, 145 (1983).
 - [16] T. Ido, *et al.*, Phys. Rev. Lett. **94**, 153001 (2005).
 - [17] R. W. P. Drever, *et al.*, Appl. Phys. B, **31**, 97 (1983).
 - [18] R. Grimm and J. Mlynek, Phys. Rev. Lett. **63**, 232 (1989); R. Grimm and J. Mlynek, Phys. Rev. A **42**, 2890 (1990).
 - [19] This coefficient is 10 times less than that measured with cold atoms[16].
 - [20] M. Petersen, *et al.*, Phys. Rev. Lett. **101**, 183004 (2008).
 - [21] T. M. Fortier, *et al.*, Phys. Rev. Lett. **97**, 163905 (2006).
 - [22] C. W. Oates, *et al.*, Phys. Rev. A **71** 023404 (2005).
 - [23] M. Baranger, Phys. Rev. **111**, 494 (1958).
 - [24] A. Bielski, *et al.*, Acta. Phys. Pol. A **90**, 1155 (1997); **105**, 217 (2004).
 - [25] W. R. Hindmarsh, *et al.*, Proc. R. Soc. London **A197**, 296 (1967).
 - [26] A. Bielski and J. Wasilewski, Z. Naturforsch. **35a**, 1112 (1980).
 - [27] J. Mitroy and J.-Y. Zhang, J. Chem. Phys. **128**, 134305 (2008) and private communication.
 - [28] P. R. Berman, *et al.*, Phys. Rev. A **20**, 2389 (1979).
 - [29] P. S. Julienne and F. H. Mies, Phys. Rev. A **34**, 3792 (1986); P. J. Leo, *et al.*, J. Phys. B **28**, 591 (1995); T.

- Orlikowski, Eur. Phys. J. D **28**, 187 (2004).
- [30] M. Takamoto, *et al.*, Nature **435**, 321 (2005).